

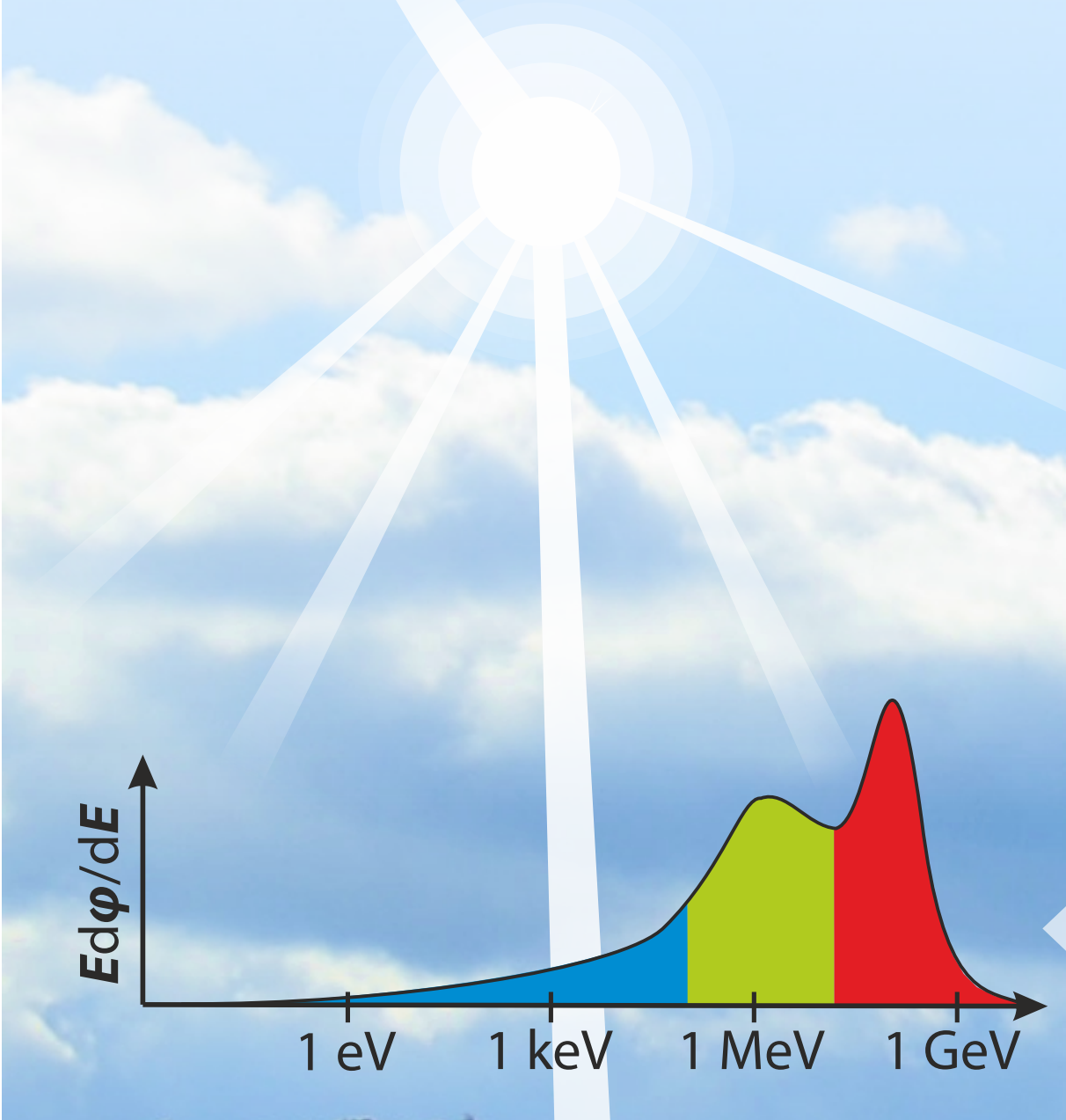
Footprint Characteristics of Cosmic-Ray Neutron Sensors for Soil Moisture Monitoring

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Software: Neutron transport was modeled with the adapted Monte-Carlo code „URANOS“, developed by the Physics Institute, Uni Heidelberg. Several tests demonstrated consistency with MCNP.
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Cosmic-Ray Neutrons

At the top of the atmosphere, primary cosmic rays generate **high-energy** neutrons, which propagate down and get moderated by air.

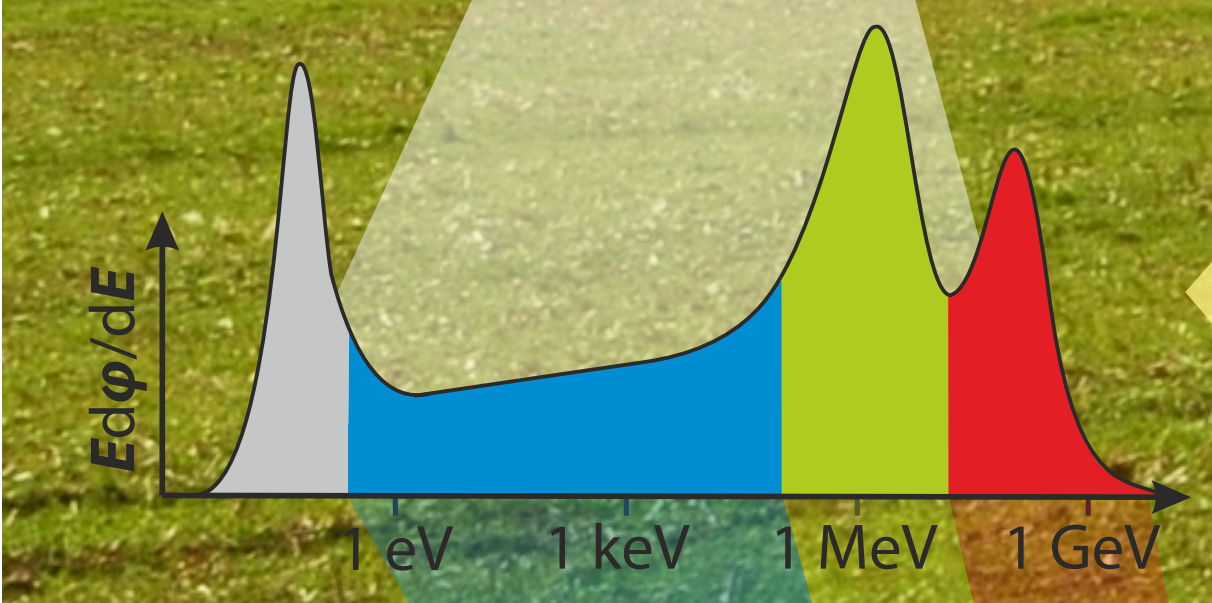
Background: The presence of hydrogen near/in the ground reduces the neutron abundance in a predictable way. Thus, area-average soil moisture can be monitored by detecting cosmic-ray neutrons in air. **Motivation:** The spatial extent of the method was previously estimated to have a 300m radius by using a simplified numerical model that released neutrons in the soil. Recent experiments, however, indicated that the sensor is extraordinarily sensitive to the near field. **Novelty:** We release simulated cosmic-ray neutrons in the atmosphere from a validated energy spectrum. A newly developed software allows to answer open questions on the horizontal weighting, penetration depth and terrain effects. **Consequences:** The footprint is smaller and not constant, which will impact data interpretation and applications.

Sensitivity to Water

High-energy neutrons are comparatively insensitive to water. At lower energies, particularly in the **blue** domain, hydrogen can effectively moderate neutrons. **Thermal** neutrons are slow and sensitive also to other chemical components.

New Model

The **full** validated spectrum (from Sato & Niita, RR, 2006) is released above ground.



Detection

A moderated ³He detector counts neutrons of the **low-energy** domain.

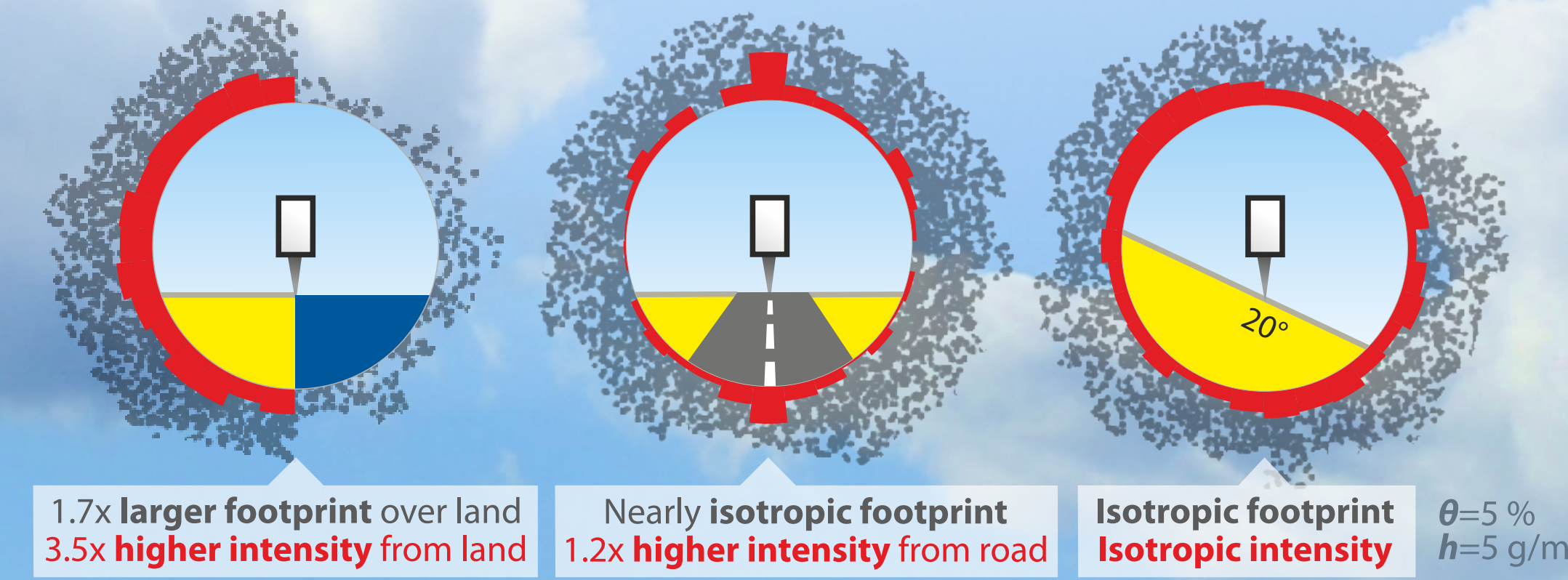
dry soil, $\theta = 1\%$
 wet soil, $\theta = 50\%$

Radial Sensitivity

The average soil moisture estimate from neutrons is neither an equally nor an exponentially weighted mean. Most of the detected neutrons **N** probed the soil within the very first meters around a sensor. The **horizontal weighting** rather is:

$$dN(r, \theta, h)/dr \approx \begin{cases} F_1 e^{-F_2 r} + F_3 e^{-F_4 r}, & r \leq 50 \text{ m} \\ F_5 e^{-F_6 r} + F_7 e^{-F_8 r}, & r > 50 \text{ m} \end{cases}$$

where $F_i(\theta, h)$

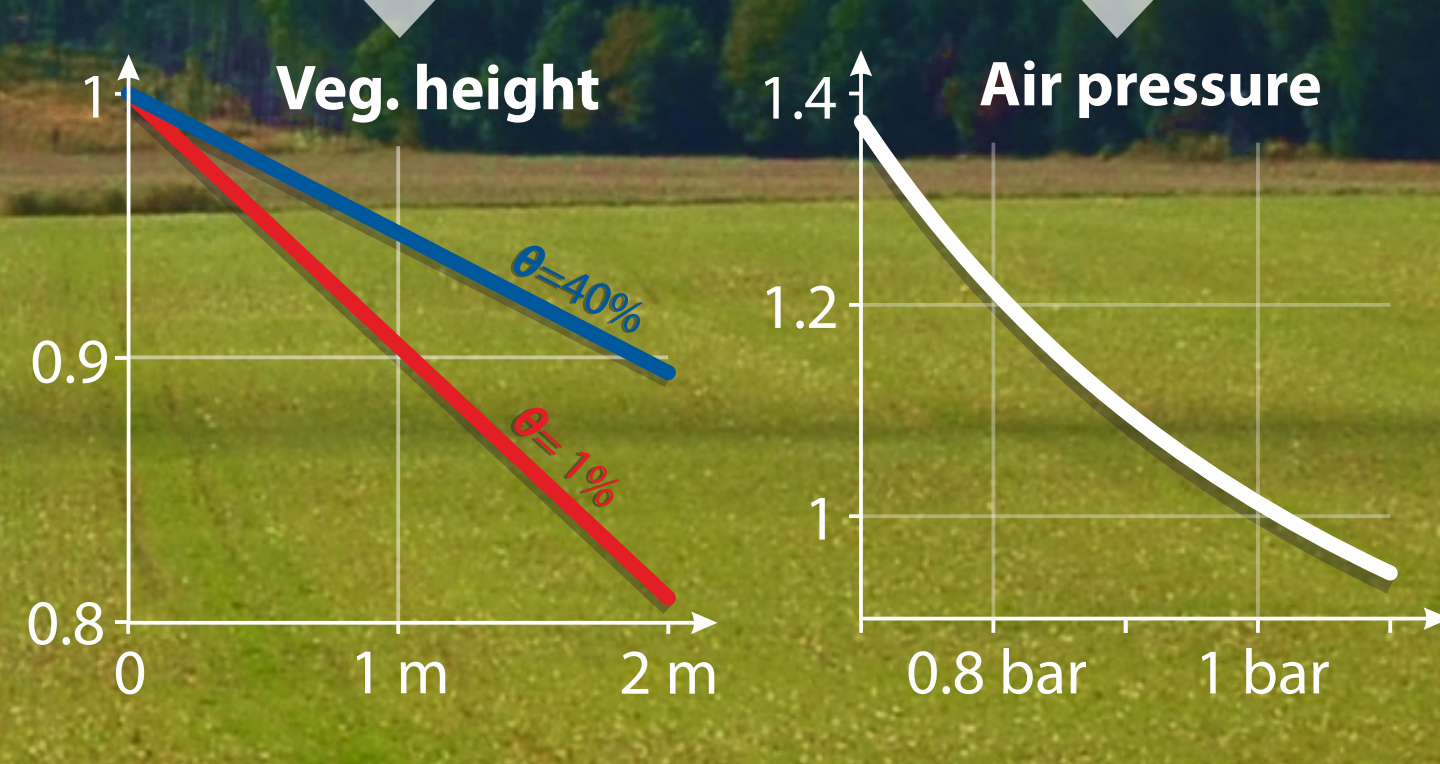


Anisotropy & Terrain Effect

From which direction and distance do neutrons preferably originate (grey)? What is their directional intensity (red)?

Scaling (veg, p)

The footprint size further varies **by a factor** due to vegetation or altitude.



Mixing in Air

Neutrons are able to travel hundreds of meters from origin (first contact with the soil) to detection.

Footprint Radius (θ, h)

The area under the curve represents the total detected neutron intensity. The footprint R_{86} is the distance within which 86% of the detected neutrons originated.

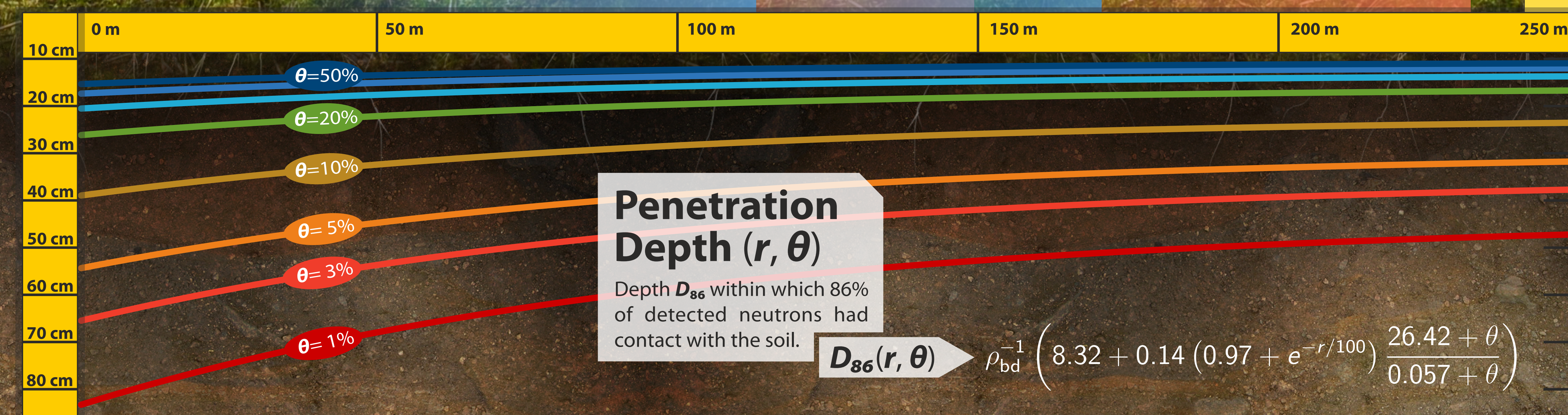
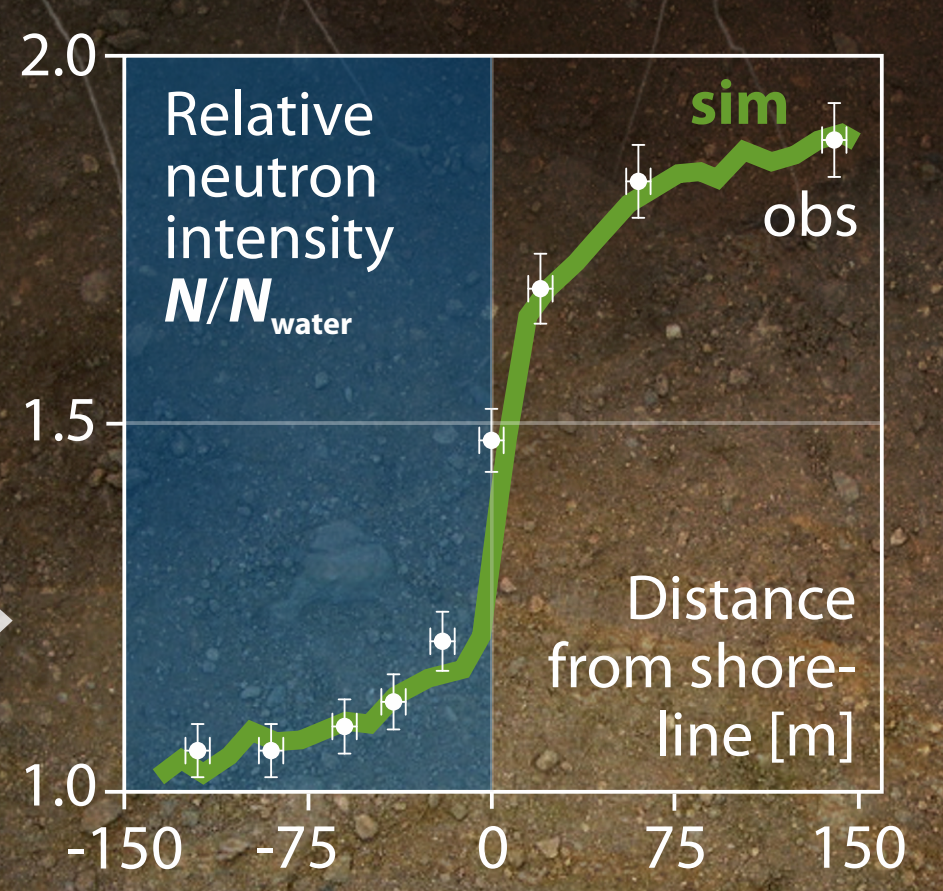
$R_{86}(\theta, h)$ 113 m 148 m 171 m 232 m

Old Model

Applies a **low-energy** source in the soil, based only on incoming **high-energy** neutrons.

Evidence

The revised model was tested by the reproduction of transect experiments across a **water-land** interface. The data also indicates the extent of the footprint and high sensitivity in the near-field.



Penetration Depth (r, θ)

Depth D_{86} within which 86% of detected neutrons had contact with the soil.

$$D_{86}(r, \theta) \rho_{bd}^{-1} \left(8.32 + 0.14 (0.97 + e^{-r/100}) \frac{26.42 + \theta}{0.057 + \theta} \right)$$

